

HAWC CDR

February 1-2, 2001

Data Processing Software

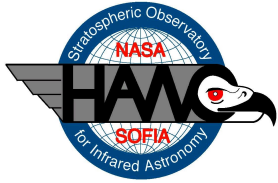
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HAWC Data-Taking Hierarchy and Terms: Stare & Nod



- **HAWC detectors: “sample-rate data”**
 - HAWC outputs images at ADC sample rate (1-4 kHz)
- **Chopping secondary: “chop-rate data”**
 - telescope secondary chops between source and reference positions at 10-20 Hz
- **Telescope nodding: “nod pair data”**
 - telescope nods to switch source & reference sky beams at < 1 Hz
 - multiple nods for each position in a dither series
 - total integration time of a few tens of sec per nod position
- **Telescope dithering: “dither series”**
 - telescope moves through a pre-defined sequence of nod positions
 - offsets are small, non-integral numbers of pixels, to smooth over detector nonuniformities
- **Telescope mapping: “map”**
 - telescope executes a sequence of dithered stare & nod observations
 - images obtained for a grid of map points
 - large separations between points (of order 80% of detector dimensions)



HAWC Observing Mode Parameters (I.e. Astronomical Observing Template [AOT])



Table 1: HAWC Mapping (with dither) AOT

Parameter	Data Type	Example script line ²	Observing Step (see above)
Observing mode	Mode	DITHER OBSMODE set	1
Source position (RA)	(h, m, s, .ss)	23 34 45 00 S_RA set	1
Source position (dec)	char (north/south=+/-) integer (d, m, s, .ss)	+ 19 20 30 00 S_DEC set	1
Dither ³	Pattern	1 DITHER set	7
Mapping offset (RA)	floating point, (arcsec)	175.0 M_RA set	1
Mapping offset (dec)	floating point, (arcsec)	175.0 M_DEC set	1
# Map Columns	integer	3 M_COL set	1
# Map Rows	integer	3 M_ROW set	1
Chop amplitude	floating point, (arcsec)	20.0 CHOP_RA_1 set	1
Chop angle	floating point, (degrees)	0.0 CHOP_ANGLE set	1
Chop mode	word	2POINT CHOP_MODE set	1
Aperture	integer, (selection)	1 APER set	2
Filter (LWC)	integer, (selection)	1 FSWC set	3
Chop cycles per frame	integer	10 CPF set	4
Frames per nod (<i>n</i>)	integer	10 FPN set	5
Integration time	floating point, (sec)	1.0 ITIME set	5
# Nod Pairs/position	Integer	2 NPAIRS set	6
Required Source S/N	integer	50 SSN set	11
Required Fluxcal S/N	integer	100 FSN set	11
Estimated Time of Observation ³	Integer		



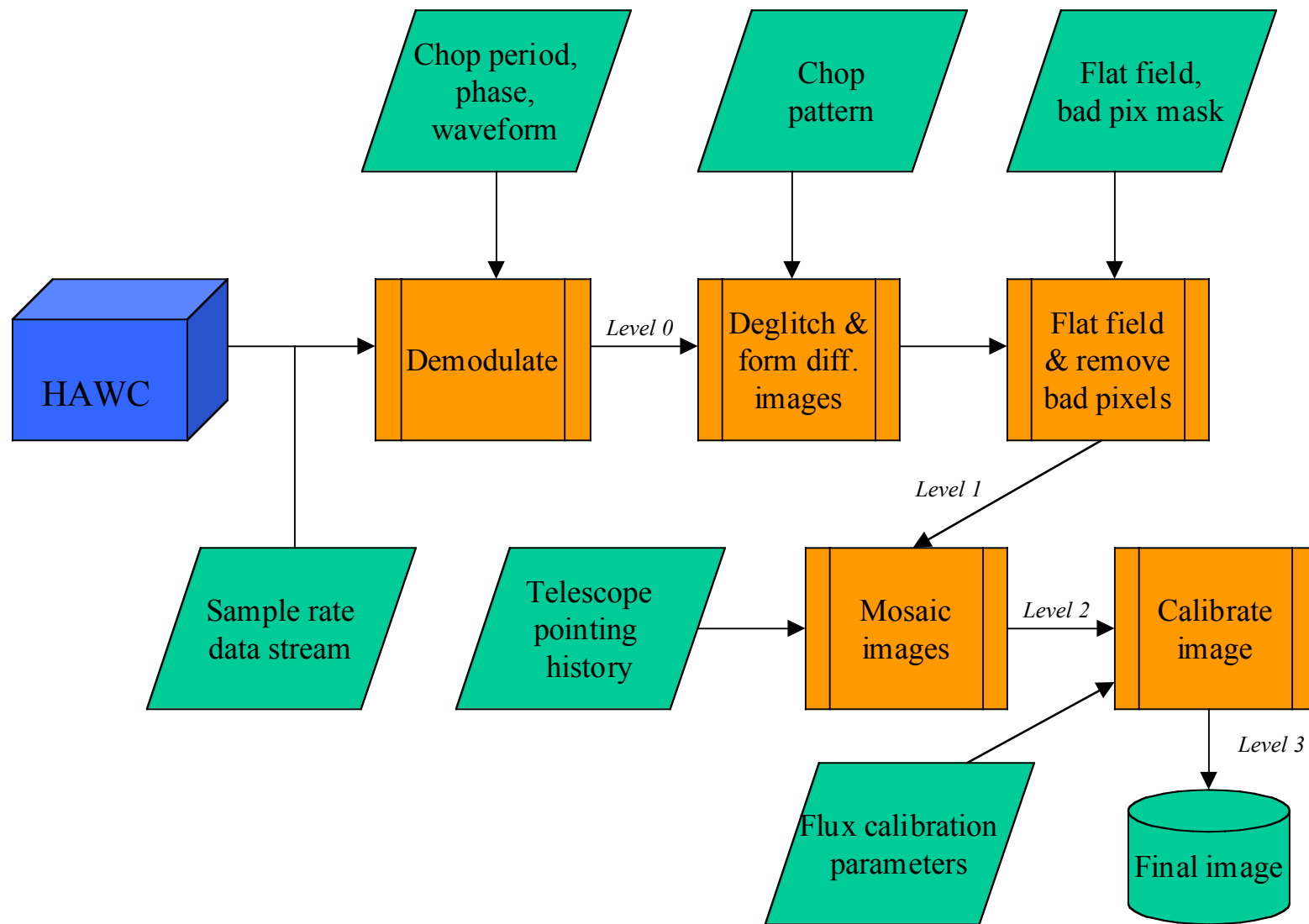
HAWC Data Processing Pipeline



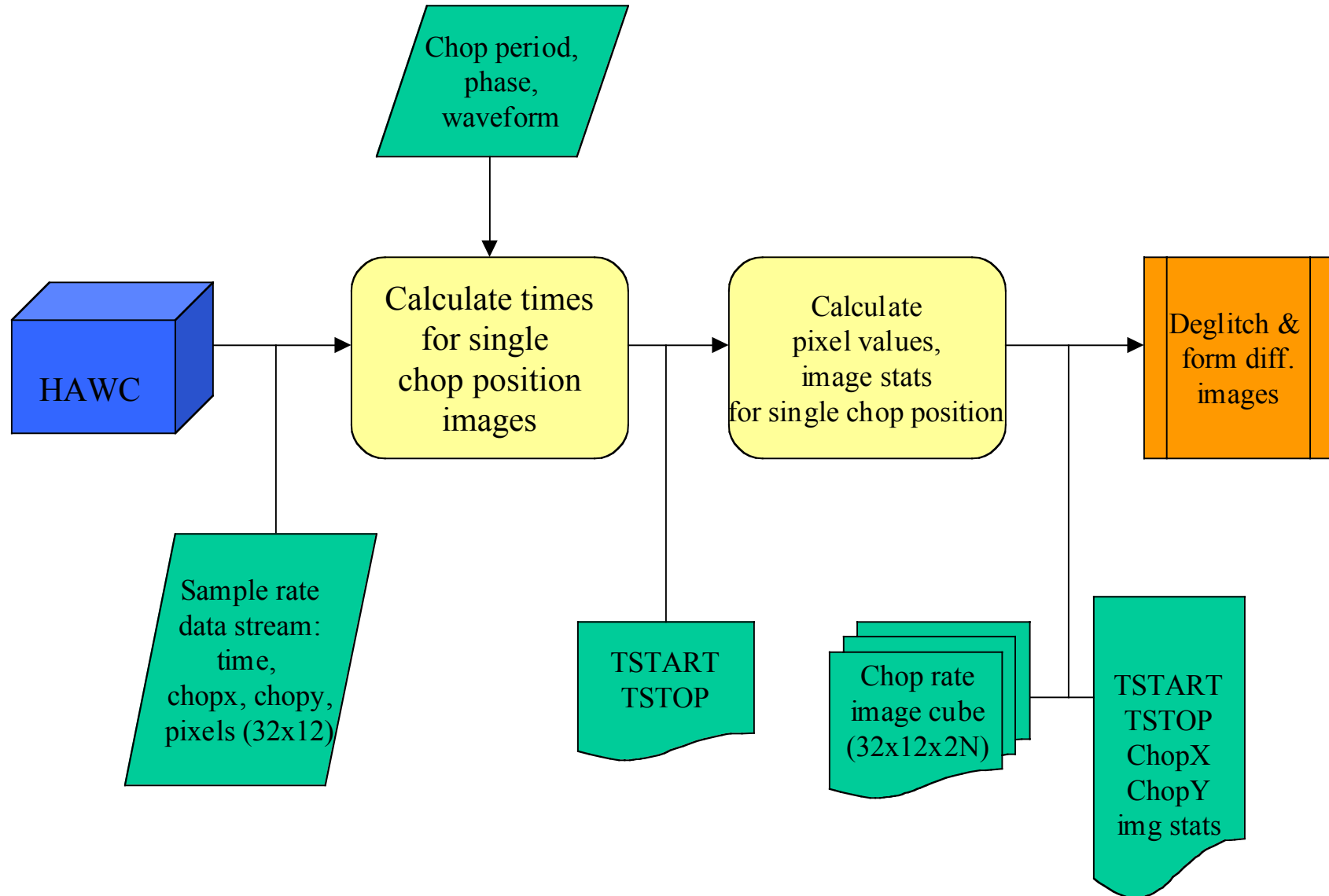
-
- **Demodulate sample-rate (ADC rate) data to produce chop-rate image data**
 - **Deglitch chop-rate data and generate chop-rate difference images**
 - **Co-add chop-rate difference images for each telescope pointing in a “stare & nod” sequence**
 - **Process coadded chop-rate difference images to remove detector artifacts & nonuniformities**
 - **Form nod pair difference image for each position of a dithered series of telescope pointings**
 - **Combine dithered image series into a mosaic & establish image coordinate system**
 - **Combine dithered mosaics obtained over map grid into a map image**
 - **Calibrate map image pixel intensities**

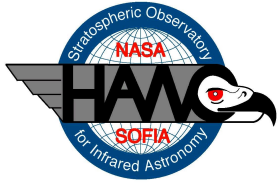


HAWC Data Pipeline: Stare & Nod Mode



Demodulation Process

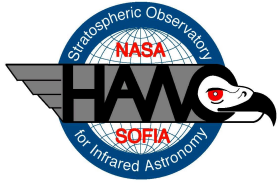




Pipeline Data Products



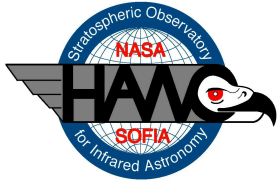
- Level 0: Chop-rate image cube**, a $32 \times 12 \times 2N$ image (where N is the number of chop cycles per nod position)
- Level 1: Nod pair difference images**, integrated over M nod pairs (where M is large enough to cover substantial integration time)
- Level 2: Mosaiced image(s)**, result of dithered set of stare/nod observations at one or more map positions
- Level 3: Calibrated map image**, result of ensemble of dithered images taken over map grid (flux-calibrated, with world coordinate system specified)



Pipeline Algorithm Development: Approach



- **Modular, extensible**
 - Pipeline as a set of stand-alone tools w/ well-defined inputs & outputs
 - isolation of I/O from data processing
- **Emphasis on rapid prototyping of algorithms**
 - IDL as prototyping language
 - migration to other languages (e.g., C or C++) where necessary (speed, memory considerations, e.g.)
- **reuse/adapt existing algorithm designs, code where feasible**
 - draw on extensive KAO experience (e.g., for deglitching schemes)
 - draw on experience with South Pole IR data system
 - investigating overlap with FORCAST pipelines
- **Use simulation tools to test & refine algorithms**



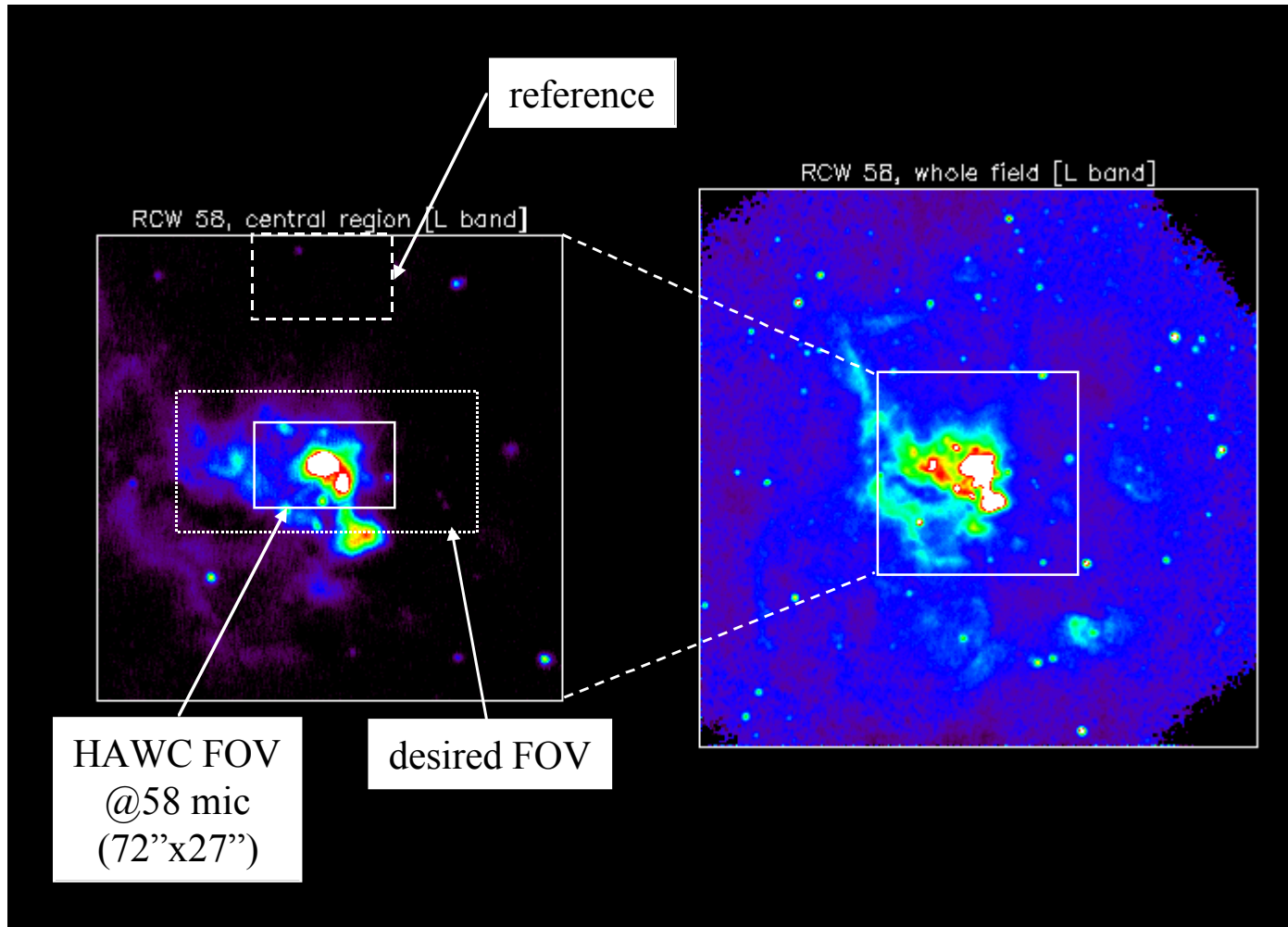
Pipeline Algorithm Development: Status



- **Simulator under development (in IDL)**
- **Prototype demodulator written**
 - prototyped in IDL; migrated to C to optimize speed & portability
- **Prototype image generator written**
 - makes use of algorithms from SPIREX/Abu (near-IR image) processing pipeline
- **Prototypes successfully integrated into the DCS**
 - integrated simulator and pipeline demonstrated at DCS Prototype Review (Nov. '00)



Example: Processing data from the HAWC Simulator





HAWC Simulator AOT GUI



HAWC Parameters

	value	Unit
Nsamples	6000	N/A
Sfrequency	1000.00	Hz
Cfrequency	20.0000	Hz
BandType(1-4)	1	N/A
Aperture	1	N/A

SET_HAWC

Telescope

	value	unit
NodTime	1.50000	second
NodPair	4	N/A
ChopAmp	90.0000	arcsec
ChopAngle	0.0000	degree

dither status: single image

SET_TSCP DITHER MOSAIC
 DONE

Mosaic Pattern

Position/Offset Selection

5 pnts
 9 pnts
 UserDfn

	value	unit
Size/StepRA	45.0	arcsec
Size/StepDEC	20.0	arcsec

OK RESET

Dither Pattern

Position/Offset Selection

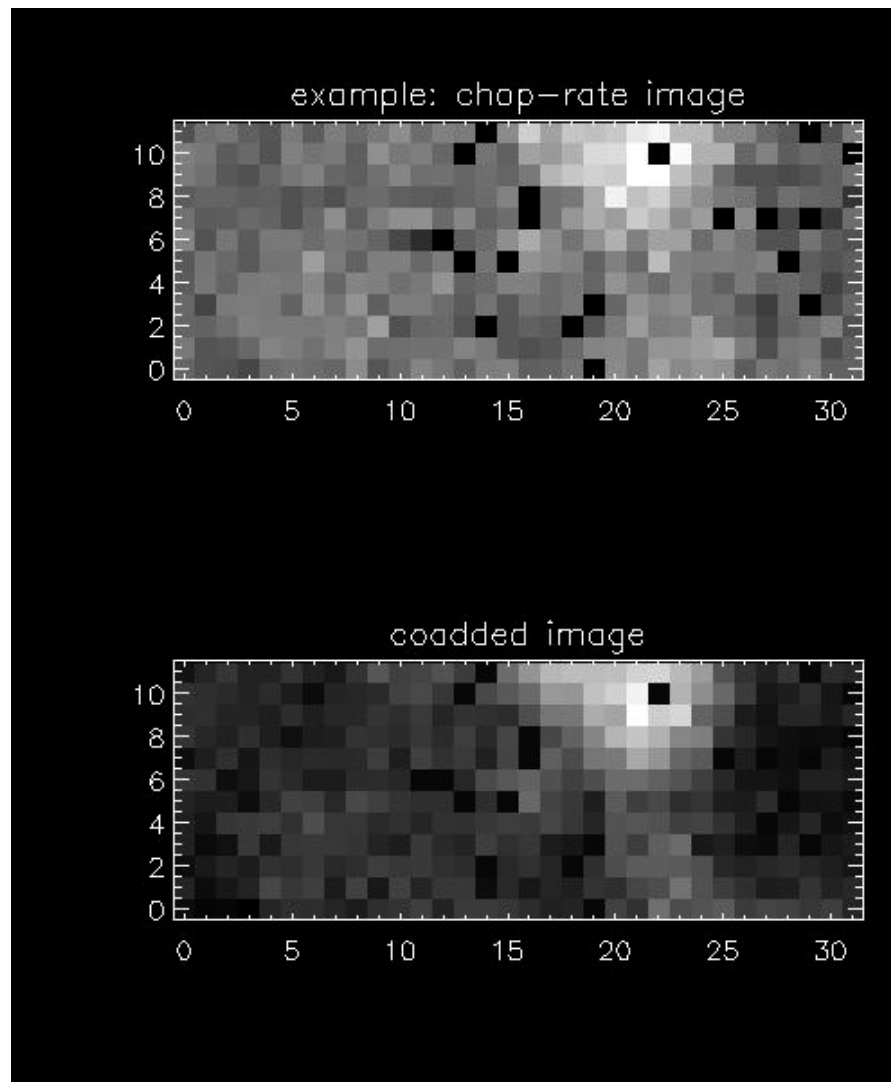
5 pnts
 9 pnts
 UserDfn

	value	unit
Size/StepRA	9.0	arcsec
Size/StepDEC	9.0	arcsec

OK RESET



Level 0 & 1 output from the prototype pipeline

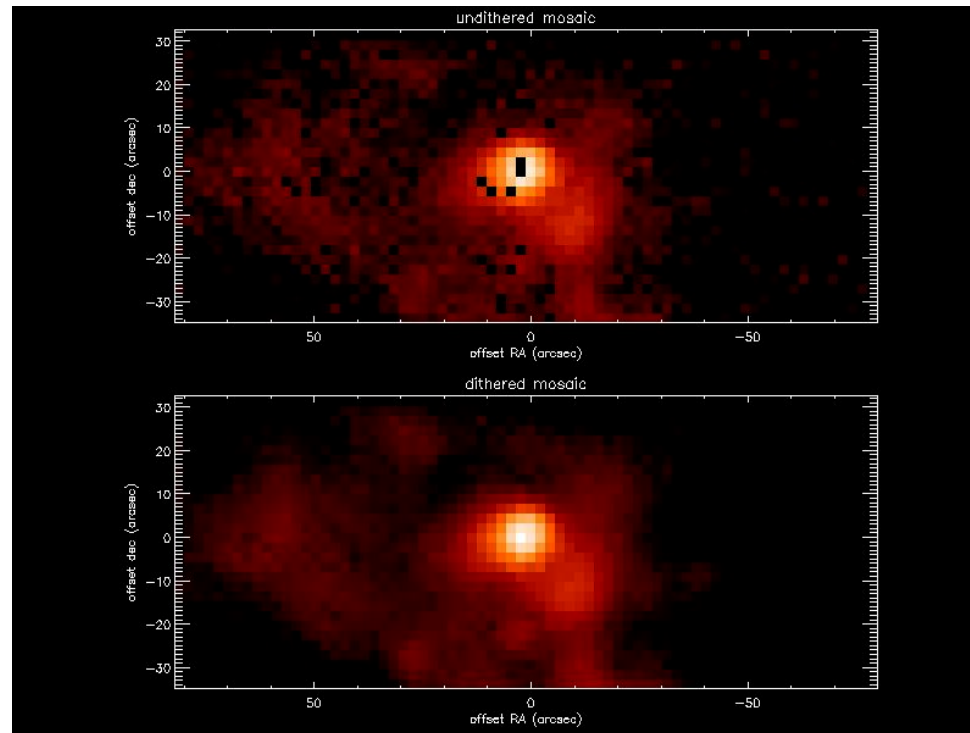


*Top: difference
of Level 0 (chop-
rate) images*

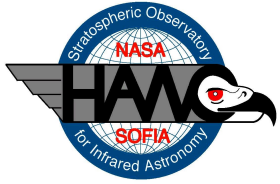
*Bottom: Level 1
(nod pair
difference) image*



Level 2 output from the prototype pipeline



Shown: comparison of mosaics constructed from individual, coadded images (top) vs. dithered sets of coadded images (bottom)



The Level 3 HAWC data product: *a flux-calibrated far-infrared image*



- **Calibration models and algorithms under development at Yerkes**
 - as part of our filter optimization work we have already developed algorithms which take a source spectrum and determine the W/pixel at the detector after the source has passed through a model atmosphere, telescope and HAWC optics
- **Pipeline will incorporate calibration algorithms**
 - Yerkes models to serve as basis for algorithms
 - algorithms (code) & parameters (calibration data files) kept separate to enable improvements in data products as calibration knowledge improves
- **Simulator will facilitate calibration algorithm testing**
 - supports introduction of incrementally improving source, atmosphere, and instrument models